

## Organometallic Photochemistry in Supercritical Fluids: the Reaction of H<sub>2</sub> with [( $\eta^5$ -C<sub>5</sub>H<sub>5</sub>)M(CO)<sub>3</sub>] (M = Re and Mn) and the Formation of a 'Non-classical' Dihydrogen Complex of Manganese(I)

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U.v. photolysis of [(cp)M(CO)<sub>3</sub>] (cp =  $\eta^5$ -C<sub>5</sub>H<sub>5</sub>, M = Re and Mn) at room temperature in supercritical Xe pressurized with H<sub>2</sub>, leads to formation of the dihydride, [(Cp)Re(CO)<sub>2</sub>H<sub>2</sub>], and the 'non-classical' dihydrogen complex, [(cp)Mn(CO)<sub>2</sub>(H<sub>2</sub>)], which is stabilized by the high pressure of H<sub>2</sub>.

Photolysis of organometallics in the presence of H<sub>2</sub> is a convenient and fairly general method for the generation of thermally labile dihydrogen and dihydride complexes.<sup>1</sup> Unfortunately, the route is limited by the rather low solubility of H<sub>2</sub> in most common hydrocarbon solvents at room temperature. This limitation can be removed completely by carrying out the reactions in supercritical fluids, such as CO<sub>2</sub> or Xe.<sup>†</sup> These fluids are totally miscible with gaseous H<sub>2</sub> or N<sub>2</sub> yet, like hydrocarbon solvents, they can dissolve a wide range of neutral organometallic compounds.<sup>2,3</sup> Thus, for a given pressure of H<sub>2</sub>, the supercritical environment will provide a much higher concentration of 'dissolved' H<sub>2</sub> than in a conventional solvent.<sup>‡</sup>

<sup>†</sup> The critical points of these substances are CO<sub>2</sub> (31 °C, 74 atm) and Xe (17 °C, 59 atm) (10.1 atm = *ca.* 1 MPa).

<sup>‡</sup> In addition, supercritical Xe is totally transparent throughout the u.v., visible, and mid-i.r. regions, making it particularly easy to monitor reactions spectroscopically.

We have recently developed<sup>3</sup> a miniature high pressure cell§ for performing supercritical reactions at temperatures up to 100 °C. We illustrate its use with the reaction of H<sub>2</sub> and [(cp)M(CO)<sub>3</sub>] (cp =  $\eta^5$ -C<sub>5</sub>H<sub>5</sub>, M = Re and Mn). This reaction is of considerable interest because the electronic properties<sup>4</sup> of the (cp)Mn(CO)<sub>2</sub> moiety suggest that the reaction product should be a relatively stable 'non-classical' dihydrogen complex.

§ The stainless steel cell has a cold finger for condensing in the Xe. It has CaF<sub>2</sub> windows (1 cm thick) and, in these experiments, a 1.8 mm pathlength. In a typical experiment the cell contained *ca.* 3 g of Xe, 1–2 mg of [(cp)M(CO)<sub>3</sub>] and 100 atm overpressure of H<sub>2</sub> gas, giving *ca.* 175 atm total pressure at +25 °C. Under these conditions, the concentration of dissolved [(cp)M(CO)<sub>3</sub>] is less than 1/10 of the saturated solution and so no undissolved solid is present in the cell. A Cermax 300 W Xe lamp filtered to give u.v. light, 200–400 nm, was used for photolysis.

*Safety note:* experiments in supercritical fluids involve relatively high pressures of gases and should be approached with caution.

**Table 1.** Wavenumbers<sup>a</sup> cm<sup>-1</sup>, of  $\nu(\text{C-O})$  i.r. bands in supercritical Xe at 25 °C.

[(cp)Mn(CO) <sub>3</sub> ] <sup>b</sup>		[(cp)Mn(CO) <sub>2</sub> (H <sub>2</sub> )]		[(cp)Mn(CO) <sub>2</sub> (N <sub>2</sub> ) <sup>c</sup>	
2032.8	a <sub>1</sub>	1991.9 (1992.3) <sup>e</sup>	a'	2171.3	$\nu(\text{N-N})^d$
1952.3	e	1932.5 (1931.4) <sup>e</sup>	a''	1985.0	a'
2021.5	$\nu(^{13}\text{C-O})$	— <sup>f</sup>	$\nu(^{13}\text{C-O})$	1934.0	a''
1917.3	$\nu(^{13}\text{C-O})$	1903.8 (1901.5) <sup>e</sup>	$\nu(^{13}\text{C-O})$	— <sup>f</sup>	$\nu(^{13}\text{C-O})$
				1904.3	$\nu(^{13}\text{C-O})$
[(cp)Re(CO) <sub>3</sub> ]		[(cp)Re(CO) <sub>2</sub> H <sub>2</sub> ] <sup>g</sup>		[(cp)Re(CO) <sub>2</sub> (N <sub>2</sub> ) <sup>h</sup>	
2035.0	a <sub>1</sub>	2028.4	a <sub>1</sub>	2146.8	$\nu(\text{N-N})$
1946.9	e	1962.0	b <sub>1</sub>	1978.4	a'
2023.9	$\nu(^{13}\text{C-O})$	—		1925.8	a''
1012.1	$\nu(^{13}\text{C-O})$	—		—	

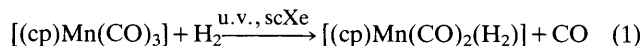
<sup>a</sup> Spectra recorded on Nicolet 730 Fourier transform (F.t.) i.r. interferometer (16K data points, 32K transform points) error  $\pm 0.2$  cm<sup>-1</sup>.

<sup>b</sup> ca. 40 other i.r. bands (fundamentals, overtones *etc.*) due to [(cp)Mn(CO)<sub>3</sub>] could be observed in the range 4400–1100 cm<sup>-1</sup>.

<sup>c</sup> Reported values,<sup>6</sup> 2169, 1980 and 1923 cm<sup>-1</sup> in hexane. <sup>d</sup> 2137.2 cm<sup>-1</sup>  $\nu(^{14}\text{N}-^{15}\text{N})$  satellite, natural isotopic abundance. <sup>e</sup> Values in parentheses for [(cp)Mn(CO)<sub>2</sub>(D<sub>2</sub>)]. <sup>f</sup> These  $\nu(^{13}\text{C-O})$  satellites obscured by other bands. <sup>g</sup> Reported values (ref. 5), 2022 and 1954 in n-heptane. <sup>h</sup> Reported values (ref. 12), 2141, 1970, and 1915 cm<sup>-1</sup> in hexane.

Figure 1 shows part of the i.r. spectra recorded after u.v. photolysis of [(cp)Re(CO)<sub>3</sub>] and [(cp)Mn(CO)<sub>3</sub>] with H<sub>2</sub> in supercritical Xe (scXe) at room temperature.<sup>¶</sup> The i.r. data are summarized in Table 1. The following points are clear from these spectra. (i) The i.r. spectra of the starting materials, [(cp)Mn(CO)<sub>3</sub>] and [(cp)Re(CO)<sub>3</sub>], are almost identical. (ii) The  $\nu(\text{C-O})$  bands of the photoproducts (coloured black in Figure 1) are quite different for the two metals. The bands of the Re species appear *between* those of [(cp)Re(CO)<sub>3</sub>], while the bands of the Mn photoproduct are shifted *down* in wavenumber relative to [(cp)Mn(CO)<sub>3</sub>]. (iii) The Re photoproduct can be identified from its  $\nu(\text{C-O})$  spectrum as the known<sup>5</sup> dihydride, [(cp)Re(CO)<sub>2</sub>H<sub>2</sub>], formed by oxidative addition of H<sub>2</sub> to the Re centre. This change in oxidation state is underlined by the fact that the  $\nu(\text{C-O})$  bands of [(cp)Re(CO)<sub>2</sub>H<sub>2</sub>] occur at substantially higher wavenumber than the corresponding bands of [(cp)Re(CO)<sub>2</sub>(N<sub>2</sub>)] (average shift +43 cm<sup>-1</sup>, see Table 1). (iv) By contrast, the  $\nu(\text{C-O})$  bands of the Mn photoproduct are very close in wavenumber to those of [(cp)Mn(CO)<sub>2</sub>(N<sub>2</sub>)] (average shift +5 cm<sup>-1</sup>), suggesting that Mn has the same oxidation state in both compounds. The photoproduct clearly contains hydrogen co-ordinated to the metal, because there is a small but definite increase (+1.5 cm<sup>-1</sup>) in the separation of the two  $\nu(\text{C-O})$  bands when D<sub>2</sub> is used instead of H<sub>2</sub>.<sup>||</sup>

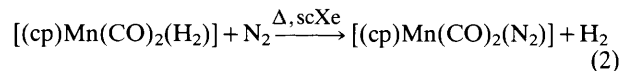
Thus, the evidence indicates that photolysis of [(cp)Mn(CO)<sub>3</sub>] in the presence of H<sub>2</sub> leads to the formation of a dihydrogen complex rather than a dihydride, equation (1).



<sup>¶</sup> For both metals, photolysis also generates free CO which, under these conditions, gives rise to two broad i.r. absorptions (*P* and *R* branches, centred at ca. 2150 cm<sup>-1</sup>, with unresolved rotational fine structure).

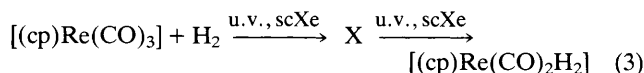
<sup>||</sup> This shift in  $\nu(\text{C-O})$  is not due to a solvent effect because the bands of [(cp)Mn(CO)<sub>3</sub>] are unshifted when D<sub>2</sub> is used instead of H<sub>2</sub>. Despite the i.r. transparency of scXe, we were unable to observe any bands directly associated with co-ordinated  $\eta^2\text{-H}_2$ , probably because there are relatively strong  $\nu(\text{C-H})$  bands due to the cp group in the region ca. 3000–2700 cm<sup>-1</sup> where  $\nu(\text{H-H})$  bands might be expected.<sup>1c</sup> Neither did we observe bands due to co-ordinated D<sub>2</sub> but the shifts in the  $\nu(\text{C-O})$  bands (see Table 1) suggest that the  $\nu(\text{D-D})$  may be masked by the much more intense  $\nu(\text{C-O})$  bands.

[(cp)Mn(CO)<sub>2</sub>(H<sub>2</sub>)] is surprisingly stable in scXe at room temperature under a pressure of H<sub>2</sub>, decaying by less than 50% overnight. If, however, the H<sub>2</sub> is vented and the cell refilled with N<sub>2</sub>, the compound reacts smoothly with N<sub>2</sub> to form the known compound<sup>6</sup> [(cp)Mn(CO)<sub>2</sub>(N<sub>2</sub>)]. Under the conditions used this exchange reaction was complete in less than 70 min,<sup>††</sup> see equation (2) and Figure 2.



Thermal reaction with N<sub>2</sub> is characteristic of dihydrogen complexes<sup>7</sup> and our observation underlines the non-classical nature of [(cp)Mn(CO)<sub>2</sub>(H<sub>2</sub>)], apparently the first Mn compound of this type.<sup>‡‡</sup> In contrast to this behaviour, the dihydride, [(cp)Re(CO)<sub>2</sub>H<sub>2</sub>], does not react thermally<sup>§§</sup> with N<sub>2</sub>, even over periods of many hours in the presence of high pressures of N<sub>2</sub> (*e.g.* 100 atm).

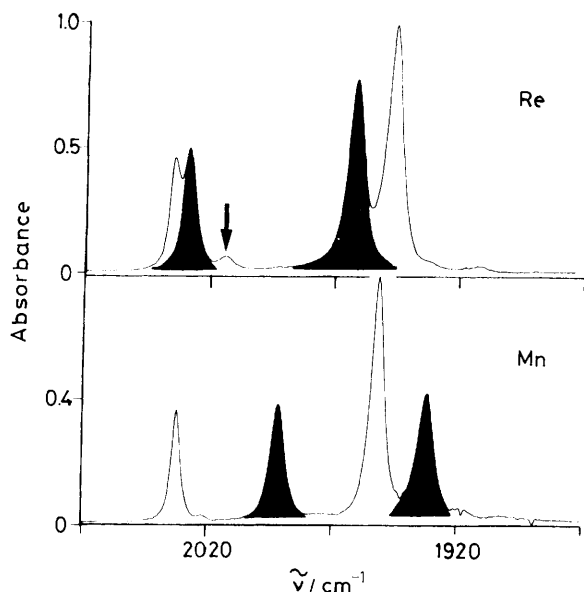
Complex [(cp)Re(CO)<sub>2</sub>H<sub>2</sub>] is not formed directly from [(cp)Re(CO)<sub>3</sub>]. There is initial formation of an unstable photoproduct, X, which decays in the dark without formation of [(cp)Re(CO)<sub>2</sub>H<sub>2</sub>]. Under u.v. irradiation, X rapidly reaches a low steady state concentration and [(cp)Re(CO)<sub>2</sub>H<sub>2</sub>] is formed. This suggests that X is an intermediate in the formation of [(cp)Re(CO)<sub>2</sub>H<sub>2</sub>].



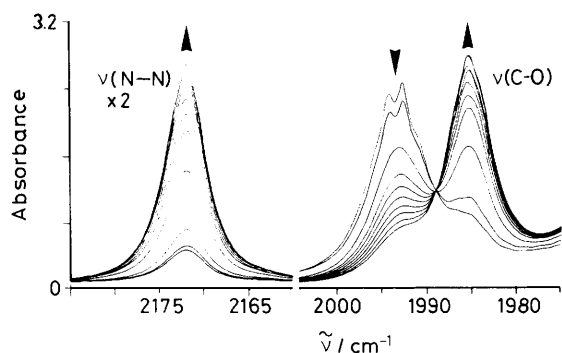
<sup>††</sup> [(cp)Mn(CO)<sub>2</sub>(N<sub>2</sub>)] can also be generated directly by u.v. photolysis of [(cp)Mn(CO)<sub>3</sub>] in the presence of N<sub>2</sub>.

<sup>‡‡</sup> Allowing for a solvent shift in the  $\nu(\text{C-O})$  bands, the [(cp)Mn(CO)<sub>2</sub>(H<sub>2</sub>)] complex generated in scXe appears to be the same as the unstable species previously observed<sup>8</sup> but not isolated on protonation of the anion [(cp)Mn(CO)<sub>2</sub>]<sup>2-</sup> in solution in tetrahydrofuran. However, this protonation experiment did not lead to any discussion on the nature of the co-ordinated H<sub>2</sub>.

<sup>§§</sup> There is, however, a photochemical reaction between [(cp)Re(CO)<sub>2</sub>H<sub>2</sub>] and N<sub>2</sub> which produces a new  $\nu(\text{N-N})$  band at 2178.4 cm<sup>-1</sup>, presumably due to a mixed dinitrogen/dihydride complex.



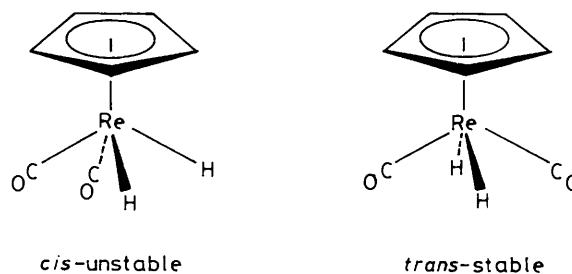
**Figure 1.** I.r. spectra in the  $\nu(\text{C-O})$  region, obtained after (a) 10 min u.v. photolysis of  $[(\text{cp})\text{Re}(\text{CO})_3]$  and (b) 5 s u.v. photolysis of  $[(\text{cp})\text{Mn}(\text{CO})_3]$  in  $\text{scXe}$  and  $\text{H}_2$  § at  $25^\circ\text{C}$ . The bands are labelled as follows: coloured, major photoproducts,  $[(\text{cp})\text{Re}(\text{CO})_2\text{H}_2]$  and  $[(\text{cp})\text{Mn}(\text{CO})_2(\text{H}_2)]$ ; Unlabelled, unreacted  $[(\text{cp})\text{M}(\text{CO})_3]$ ; arrowed, primary photoproduct which rapidly reaches steady state, see text. ¶¶



**Figure 2.** I.r. spectra in the  $\nu(\text{N-N})$  and part of the  $\nu(\text{C-O})$  regions showing the thermal reaction of  $[(\text{cp})\text{Mn}(\text{CO})_2(\text{H}_2)]$  with  $\text{N}_2$  in  $\text{scXe}$  at  $25^\circ\text{C}$ .  $[(\text{cp})\text{Mn}(\text{CO})_2(\text{H}_2)]$  was initially generated photochemically. The contents of the high pressure cell were then condensed into the cold finger of the cell, most of the  $\text{H}_2$  was vented, the cell was repressurized with  $\text{N}_2$  (ca. 100 atm) and the cold finger was warmed up. The first spectrum shown was recorded as soon as the cold finger had warmed to room temperature and subsequent spectra were taken at 10 min intervals. The band labelled ▼ is due to  $[(\text{cp})\text{Mn}(\text{CO})_2(\text{H}_2)]$  and those labelled ▲ are due to the known compound<sup>6</sup>  $[(\text{cp})\text{Mn}(\text{CO})_2(\text{N}_2)]$ . Note that the  $\nu(\text{N-N})$  band has been plotted with a  $\times 2$  expansion in the ordinate scale.

X has a  $\nu(\text{C-O})$  band, arrowed in Figure 1, at  $2013.8\text{ cm}^{-1}$  ( $14.6\text{ cm}^{-1}$  lower than that of  $[(\text{cp})\text{Re}(\text{CO})_2\text{H}_2]$ ). By comparison with matrix isolation experiments,<sup>9</sup> we suggest that X is

¶¶ In an Ar matrix, *cis*- $[(\text{cp})\text{Re}(\text{CO})_2\text{H}_2]$  has two  $\nu(\text{C-O})$  bands; the first at  $1999\text{ cm}^{-1}$  ( $13\text{ cm}^{-1}$  lower than the corresponding band of *trans*  $[(\text{cp})\text{Re}(\text{CO})_2\text{H}_2]$ ) and the second at  $1927\text{ cm}^{-1}$ , which allowing for the solvent shift would probably be masked in our  $\text{scXe}$  experiment. The formation of the intermediate X explains at least partly why the photolysis time required to generate  $[(\text{cp})\text{Re}(\text{CO})_2\text{H}_2]$  was much longer than that needed for  $[(\text{cp})\text{Mn}(\text{CO})_2(\text{H}_2)]$ , which is apparently formed directly from  $[(\text{cp})\text{Mn}(\text{CO})_3]$ . For both metals, prolonged photolysis eventually results in complete destruction of the parent  $[(\text{cp})\text{M}(\text{CO})_3]$ .



$[(\text{cp})\text{Re}(\text{CO})_2\text{H}_2]$  with the two H atoms *cis* rather than *trans* as found in the stable isomer of  $[(\text{cp})\text{Re}(\text{CO})_2\text{H}_2]$ . ¶¶ It is presumably the high concentration of  $\text{H}_2$  in the supercritical fluid, which stabilizes X sufficiently to make the subsequent photochemical generation of  $[(\text{cp})\text{Re}(\text{CO})_2\text{H}_2]$  possible.

These experiments show that there is a clear difference in the way that the Mn and Re centres react with hydrogen; Mn forms a dihydrogen complex while Re forms a dihydride. This behaviour is reminiscent of the reactivity of these species towards  $\text{HSiR}_3$  or  $\text{HSnR}_3$  compounds; Re undergoes complete oxidative addition while for Mn the process is 'arrested'.<sup>10,11</sup> In both cases the difference is presumably due to greater electron density on the Re centre. The high 'effective' concentration of  $\text{H}_2$  in a supercritical fluid is clearly stabilizing  $[(\text{cp})\text{Mn}(\text{CO})_2(\text{H}_2)]$  at room temperature. Work is now in progress to carry out such reactions on a preparative scale.

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